Volcano monitoring from space

Pierre Briole - ENS GIFT Meeting – Vienna 2009





Summary

• 1: Volcanoes on Earth • 2: Volcanic hazards • 3: Volcano monitoring from the ground • 4: Volcano monitoring from space -4a: C and L-band radar (deformations) -4b: Infra-red (plume, thermal) -4c: Visible (thermal, mapping) -4d: Ultra-violet (plume: sulfur dioxide)

Active volcanoes on Earth



More than 1500 volcanoes on the Earth are potentially active. About 70 are presently erupting.

10% of the world population is living in areas threatened by volcanoes, without considering the effects of eruptions on climate or air-traffic.

Volcanoes and the security of air traffic



Routes of some of the 100000 flight per year in the Alaska-Aleoutians area

2: Volcanic hazards 3: Volcano monitoring from the ground 4: Volcano monitoring from space

Lava flows



The large 1669 eruption of Etna

Lava flows produce irreversible damages to the land. In general the velocity of lava flow is slow and the population can escape. The flow extension depends on slope, effusion rate, duration of the eruption

Lava domes and pyroclastic flows



Eruptions of explosive volcanoes are often preceded by the growth of a lava dome.

The eruption occur when the amount of gas becomes so high that the lava dome explodes.

Pyroclastic flows are a mix of hot gas and lava blocks. They can be very fast.

Ash fall / mud flows





Ash fall and roof collapses during the 1992 eruption of Pinatubo, Philippines © USGS



Armero destroyed by a mudflow after the melting of ice on the Nevado del Ruiz volcano in 1985 © USGS

Near Pinatubo, Philippines, 1991 © USGS



Lake Nyos, Cameroun © Smithsonian



Soufrière de Guadeloupe, 2000 © JC Komorowski



Sakurajima, 2000 © P. Briole

Volcanoes inject in the troposphere H2O, CO2, SO2, H2, CO and in lower quantities H2S, HCl, HF, He, ...

Gaz

Those gases can be responsible of acid rains, pollution of aquifers, The volcanic plumes have also an impact on the climate. Some historical eruptions have induced colder climate during years.

• 3: Volcano monitoring from the ground

• 4: Volcano monitoring from space

Seismicity

Seismic sensors are the basic tool in volcano observatories

Seismometer



Seismicity at the beginning of an eruption at Piton de la Fournaise © Piton de la Fournaise volcano observatory



Strainmeter





Ground deformations



Ground deformation at a GPS station during the beginning of the November 15, 2002 eruption © P. Briole and Piton de la Fournaise volcano observatory

New ground deformation techniques have been implemented at volcanoes in the last two decades



Automated geodimeter at Piton de la Fournaise volcano © P. Briole



GPS measurements at Piton de la Fournaise © P. Briole

Gas monitoring

SOF measurements traversing the plume of Etna, Oct 7, 15:05



Etna, 2005 – COSPEC measurements of the SO2 in the volcanic plume © INGV Catania

> Stromboli, 1985, aerosols sampling for the analysis of radon daughters short lived radio-nuclides © P. Briole



Poas, 1985, gas sampling in a fumarole © JL Cheminée



Lava flow emission monitoring





Lava sampling at Piton de la Fournaise © T. Staudacher



Doppler radar at Etna volcano © OPGC Clermont-Ferrand 4: Volcano monitoring from space

-4a: C and L-band radar (deformations)
-4b: Infra-red (plume, thermal)
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The electromagnetic spectrum



• 4a: C and L-band radar (deformations)

- 4b: Infra-red (plume, thermal)
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SAR interferometry satellites

- ERS 1 1992-2000 C-band
- ERS 2 1995-2001 C-band
- ENVISAT 2002-? C-band
- **RADARSAT 1** 1995-? C-band
- RADARSAT 2 2007-?- C-band
- ALOS 2006-? L-band
- TERRASAR X 2007-? X-band
- COSMO 2007-? X-band





ENVISAT before launch © ESA

Repeat time: 16 to 35 days

- SENTINEL 1 (GMES, 2011) – C-band



Artist vue of ALOS © JAXA

L band $\lambda = 23 \text{ cm}$



C band





Deformation due to dyke injection



July 2006 eruption of Piton de la Fournaise © OPGC Clermont-Ferrand



1995 eruption of Fernandina, Galapagos © Jonnson et al.

Large scale deformations (Etna)





4b: Infra-red (plume, thermal)
4c: Visible (thermal, mapping)
4d: Ultra-violet (plume: sulfur dioxide)

Volcano monitoring in the infra red



The MODIS Rapid Response System: http://rapidfire.sci.gsfc.nasa.gov

Band	Bandwidth [µm]	Use for volcanology
1	0.620 - 0.670	Plume RGB: 1-4-3
2	0.841 - 0.876	
3	0.459 - 0.479	Plume RGB: 1-4-3
4	0.545 - 0.565	Plume RGB: 1-4-3
5	1.230 - 1.250	
6	1.628 - 1.652	
7	2.105 - 2.155	
21	3.929 - 3.989	Hot-Spot
22	3.929 - 3.989	Hot-Spot
31	10.780 - 11.280	Plume, Hot-Spot
32	11.770 - 12.270	Plume, Hot-Spot

MODIS bands

The case of MODIS (Moderate Resolution Imaging Spectrometer)

- 2 NASA satellites: Terra (since 1999), Aqua (since 2002)
- Spatial resolution = 1km (some bands available in 250m and 500m)
- Temporal resolution = 1 day- and 1 night pass of each satellite every 48 hours = 4 observations in 2 days).
- 36 bands: visual, near-IR, mid-IR, thermal IR



True (visible) colours

MODIS on TERRA – View of Etna on 16 April 2009

Plume and deposits observed by MODIS



© Alaska Volcano Observatory

Redoubt volcano, Alaska, 16 April 2009

5 April – True colours



Redoubt Alaska 5 April 2009 (MODIS on Terra – 250m – True colours)

5 April – Composition 721



Redoubt Alaska 5 April 2009 (MODIS on Terra – 250m – composition 721) © NASA

16 April – True colours



Redoubt Alaska 16 April 2009 (MODIS on Terra – 250m – True colours)

16 April – True colors



Redoubt Alaska 16 April 2009 (MODIS on Terra – 250m – True colours)

16 April – Composition 721



Redoubt Alaska 16 April 2009 (MODIS on Terra – 250m – composition 721) © NASA

4c: Visible (thermal, mapping) 4d: Ultra-violet (plume: sulfur dioxide)

Volcano monitoring in visible band



Landsat and the Enhanced Thematic Mapper (ETM) © EOSAT



Etna, 2001 – LANDSAT 7



Merapi 2006, ASTER © NASA

Band Number	μm	Resolution
1	0.45-0.515	30 m
2	0.525-0.605	30 m
3	0.63-0.69	30 m
4	0.75-0.90	30 m
5	1.55-1.75	30 m
6	10.4-12.5	60 m
7	2.09-2.35	30 m
8	0.52-0.9	15 m

Evolution of the vegetation after a large eruption

-Pinatubo: comparison of SPOT data acquired between 1991 and 1998

- Red shows vegetation, the ash deposits are light blue

- Summit is at lower right



Near real time monitoring of hot spots

http://modis.higp.hawaii.edu/



MODIS at Fernandina 11 April 2009 © NASA

Fernandina lava flow on 19 April 2009 © NASA

• 4d: Ultra-violet (plume: sulfur dioxide)

Volcano monitoring in the ultraviolet

The Ozone Monitoring Instrument on AURA © NASA

Item	Parameter
Wavelength range:	1
Visible:	350 - 500 nm
UV:	UV-1, 270 to 314 nm, UV-2 306 to 380 nm
Spectral resolution:	1.0 - 0.45 nm FWHM
Spectral sampling:	2-3 for FWHM
Telescope FOV:	114Å (2600 km on ground)
IFOV:	3 km, binned to 13 x 24 km
Detector:	CCD: 780 x 576 (spectral x spatial) pixels
Mass:	65 kg
Duty cycle:	60 minutes on daylight side
Power:	66 watts
Data rate:	0.8 Mbps (average)



AURA satellites during its building © NASA

Monitoring of volcanic plumes



SO₂ at Etna in 2001 measured by GOME on ERS2 © DLR-ESA

SO₂ at Fernandina measured on 14 April 2009 by the OMI (Ozone Monitoring Instrument) on board of AURA © NASA

Large volcanic plumes



The two eruptions with stronger impact on the climate in the last 30 years were the 1982 El Chicon (Mexico) and the 1991 Pinatubo (Philippines) eruptions. Both could be accurately observed (SO₂ content) by Ozone Mapping Spectrometers

